The Strong Onsager Conjecture

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Purdue University

FD2 Reunion Seminar August 1st, 2023

Navier-Stokes and Euler Equations

$$\partial_t u + \operatorname{div} (u \otimes u) + \nabla p - \nu \Delta u = f$$

div $u = 0$

$$\circ \ u(t,\cdot):\mathbb{T}^3 o\mathbb{R}^3$$
, $p(t,\cdot):\mathbb{T}^3 o\mathbb{R}$, $f(t,\cdot):\mathbb{T}^3 o\mathbb{R}^3$

 \circ NSE - $\nu > 0$, Euler - $\nu = 0$

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- Focus: Turbulent regime $\nu \rightarrow 0$
- Facts: (1) Anomalous dissipation of energy, (2) ⁴/₅-law, (3) intermittency

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- Focus: Turbulent regime $\nu \rightarrow 0$
- Facts: (1) Anomalous dissipation of energy, (2) ⁴/₅-law, (3) intermittency
- Onsager program: Build solutions to the PDEs consistent with experiments and numerics!

Main Theorem

Theorem (Giri-Kwon-N., '23)

For any fixed $\beta < 1/3$, there exist weak solutions to the 3D Euler equations

$$\partial_t u + \operatorname{div} (u \otimes u) + \nabla p = 0$$

 $\operatorname{div} u = 0$

which, in addition, dissipate the total kinetic energy $1/2 ||u(t)||_{L^2}^2$, belong to $C_t^0 B_{3,\infty}^\beta(\mathbb{T}^3)$, and satisfy the local energy inequality

$$\partial_t \left(rac{1}{2} |u|^2
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- Sharpness: If $\beta > 1/3$, sol'ns in $C_t^0 B_{3,\infty}^\beta$ satisfy local energy equality (Duchon-Robert)
- $\circ~$ Previous results: $C_t^0 C^\beta$ with $\beta <$ $^1\!/_{15}$ (Isett), $\beta <$ $^1\!/_{7}$ (De Lellis-Kwon)

o Navier-Stokes equations for an incompressible fluid of constant density

$$\partial_t u + \operatorname{div} (u \otimes u) = \frac{1}{\operatorname{Re}} \Delta u - \nabla p + f$$

div $u = 0$

u is velocity, p is pressure, f is an external force

• The Reynolds number

$$\operatorname{Re} = \frac{UL}{\nu} = \frac{(\operatorname{characteristic velocity}) \cdot (\operatorname{characteristic length})}{\operatorname{kinematic viscosity}}$$

 $\circ~$ Euler equations correspond to Re = $\infty,$ or $\nu=0$

What happens as the Reynolds number increases?



Flow behind a cylinder at Re = 1.54

What happens as the Reynolds number increases?



Flow behind a cylinder at Re = 140

What happens as the Reynolds number increases?



Flow behind a grid at Re = 1800

- Homogeneous isotropic turbulence arises at large Reynolds numbers (or small ν)
- What about anomalous dissipation, the 4/5 law, and intermittency?



Contour plot of dissipation in a turbulent velocity field Source: Kaneda-Ishihara '05

$$\partial_t u^{\nu} + (u^{\nu} \cdot \nabla) u^{\nu} = \nu \Delta u^{\nu} - \nabla p^{\nu}, \qquad \text{div } u^{\nu} = 0$$

• Pointwise energy balance for smooth solutions

$$\partial_t \left(\frac{1}{2} |u^{\nu}|^2\right) + \operatorname{div} \left(\left(\frac{1}{2} |u^{\nu}|^2 + p^{\nu}\right) u^{\nu} - \nu \nabla \frac{|u^{\nu}|^2}{2} \right) = -\nu |\nabla u^{\nu}|^2$$

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 $\circ~$ Integrating in \mathbb{T}^3 and from 0 to ${\it T},$ we have

$$\frac{1}{2} \| u^{\nu}(T, \cdot) \|_{L^{2}(\mathbb{T}^{3})}^{2} - \frac{1}{2} \| u^{\nu}(0, \cdot) \|_{L^{2}(\mathbb{T}^{3})}^{2} = -\int_{0}^{T} \nu \| \nabla u^{\nu}(t, \cdot) \|_{L^{2}(\mathbb{T}^{3})}^{2} dt$$

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 $\circ~$ Thus smooth Euler solutions conserve energy, and dissipation in smooth Navier-Stokes solutions is caused by $\nu\Delta u^{\nu}$

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• Energy balance for weak solutions (obtained by mollifying and passing to the limit) $\partial_t \left(\frac{1}{2}|u^{\nu}|^2\right) + \text{div} \left(\left(\frac{1}{2}|u^{\nu}|^2 + p^{\nu}\right)u^{\nu} - \nu \nabla \frac{|u^{\nu}|^2}{2}\right) + \nu |\nabla u^{\nu}|^2 + D[u^{\nu}] = 0$

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- The nonlinearity contributes the *Duchon-Robert measure* $D[u^{\nu}](t,x) = \lim_{\ell \to 0} \frac{1}{4} \int_{\mathbb{T}^3} \nabla \phi_{\ell}(z) \cdot (u(t,x+z) - u(t,x)) |u(t,x+z) - u(t,x)|^2 dz$

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- Zeroth law of turbulence (no proof exists!)

$$\varepsilon = \liminf_{\nu \to 0} \underbrace{\left\langle \nu | \nabla u^{\nu} |^2 + D[u^{\nu}] \right\rangle}_{\varepsilon^{\nu}} > 0$$

· Caffarelli-Kohn-Nirenberg's "suitable solutions" to Navier-Stokes satisfy

$$u^{\nu} \in L^{\infty}_t L^2_x \cap L^2_t W^{1,2}_x, \qquad D[u^{\nu}] \geq 0$$

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vanishes if $u^{
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• Therefore if u dissipates energy, then u^ν cannot remain bounded in $L^3_tB^\alpha_{3,\infty,{\rm x}}$ as $\nu\to 0$ for $\alpha>{}^1\!/\!{}_3$

• K41 Assumptions: the zeroth law ($\varepsilon > 0$), translation, rotation, and scaling symmetries for law of $u^{\nu}(t, x + \ell \hat{z}) - u^{\nu}(t, x)$ (here $\ell > 0, \hat{z} \in \mathbb{S}^2$)

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- K41 Claims: longitudinal structure functions satisfy $S^{\parallel}_{p}(\ell) = \langle ((u^{\nu}(t, x + \ell \hat{z}) - u^{\nu}(t, x)) \cdot \hat{z})^{p} \rangle \approx (\varepsilon \ell)^{p/3}$

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Local, deterministic ⁴/₅ law (Eyink, '02)

$$\begin{split} \lim_{\ell \to 0} \frac{1}{\ell} S_3^{\parallel}(\ell) &= \lim_{\ell \to 0} \frac{1}{\ell} \int_{\mathbb{T}^3} \int_{\mathbb{S}^2} \left[\left(u^{\nu}(t, x + \ell \hat{z}) - u^{\nu}(t, x) \right) \cdot \hat{z} \right]^3 \, d\hat{z} \, dx \\ &= -\frac{4}{5} D[u^{\nu}] \end{split}$$

Onsager's Conjecture and the $L_t^{\infty} C_x^{1/3}$ Threshold

 "It is of some interest to note that in principle, turbulent dissipation as described could take place just as readily without the final assistance by viscosity. In the absence of viscosity, the standard proof of the conservation of energy does not apply, because the velocity field does not remain differentiable!" – Onsager, '49

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and conservation of energy follows from D[u] = 0, which holds if $u \in L^3_t B^{\alpha}_{3,\infty}$ for $\alpha > 1/3$ (Eyink '92, Constantin-E-Titi '94, Duchon-Robert '00)

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• If $\alpha < \frac{1}{3}$, the kinetic energy of 3D Euler solutions need not be conserved (Isett '18) and can dissipate (Buckmaster-De Lellis-Székelyhidi-Vicol '19)

Adding to the story: local energy inequality and intermittency

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 Conservation of energy requires only L¹_tB³_{3,∞} for α > 1/3, but dissipative solutions belong to C^α_{t,x} for α < 1/3 ... is this merely a curiosity concerning function spaces?





- Onsager, unpublished work "[Anomalous scaling for ζ₂] would require a "spotty" distribution of the regions in which the velocity varies rapidly"
- Kolmogorov '62 "I have formulated appropriate modifications to the two similarity hypotheses that I put forward in 1941 ..."
- Chen, Dhruva, Kurien, Sreenivasan, Taylor '05 "It is now believed that the scaling exponents of moments of velocity increments are anomalous ... anomalous scaling is a genuine result worth of a serious theoretical effort."
- Iyer, Sreenivasan, Yeung '20 "The ⁴/₅-ths law holds in an intermediate range of scales and the second-order exponent over the same range of scales is *anomalous*, departing from the self-similar value of ²/₃."
- See also Ishihara-Kaneda-Gotoh, Frisch, Anselmet-Gagne-Hopfinger-Antonia, ...

Takeaway: $B_{3,\infty}^{1/3} \cap L^{\infty}$ may be the correct space



- **Symmetry assumptions:** Turbulence is *isotropic*, *homogeneous*, *but not purely self-similar* ... *fewer eddies of higher intensity!*
- Dissipativity assumption: Dissipation occurs even in the absence of viscosity
- Implications for regularity: Cantor function, Heaviside function $(B_{p,\infty}^{1/p})$



Strong Onsager Conjectures

Consider weak solutions u to the Euler equations, with the local energy balance

$$\begin{cases} \partial_t u + \operatorname{div} (u \otimes u) + \nabla p = 0\\ \operatorname{div} u = 0\\ \partial_t \left(\frac{1}{2} |u|^2\right) + \operatorname{div} \left(\left(\frac{1}{2} |u|^2 + p\right) u\right) + D[u] = 0\,. \end{cases}$$

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Conjecture (Strong L³ Onsager Conjecture)

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Conjecture (Strong L^3 Onsager Conjecture in the Inviscid Limit)

There exist sequences of suitable solutions $\{u^{\nu_j}\}_{j\in\mathbb{N}}$ to Navier-Stokes with viscosity $\nu_j \to 0$ which are uniformly bounded in $L_t^{\infty} B_{3,\infty}^{\alpha}$ converging to the 3D Euler solutions described above.

Theorem (Buckmaster-Masmoudi-N.-Vicol, '21)

There exist solutions to 3D Euler which have decreasing kinetic energy and belong to $L_t^{\infty} H_x^{1/2-}$ (intermittent, but no $^4/_5$ law or local energy inequality)

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Theorem (Giri-Kwon-N., '23b)

There exist solutions to 3D Euler which have decreasing kinetic energy, belong to $L_t^{\infty} \left(L_x^{\infty-} \cap B_{3,\infty}^{^{1/3-}} \right)$, and satisfy $D[u] \ge 0$ (intermittent, $^4/_5$ law, local energy inequality)

Basic Strategy: Construct a weak solution $u = \lim_{q \to \infty} u_q$ in $L_t^{\infty} \left(L^{\infty -} \cap B_{3,\infty}^{1/3-} \right)$ via Nash iteration

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- Intermittent pressure π_q : Control u_q (and any other important functions) in terms of π_q , and propagate scaling laws comparing different terms in the wavelet expansion of π_q

Basic Strategy: Construct a weak solution $u = \lim_{q \to \infty} u_q$ in $L_t^{\infty} \left(L^{\infty -} \cap B_{3,\infty}^{1/3-} \right)$ via Nash iteration

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Two questions: How to set up the induction, and how to propagate the inductive assumptions?

Wavelet-inspired, $B^{1/3-}_{3,\infty} \cap L^{\infty-}$ inductive set-up

 $\,\circ\,$ Assume the existence of $(\mathit{u_q}, \mathit{p_q}, \mathit{R_q}, \mathit{\Phi_q}, \pi_{q})$ satisfying

$$\partial_t u_q + \operatorname{div} \left(u_q \otimes u_q \right) + \nabla p_q = \underbrace{\operatorname{div}_{\times} \left(R_q - \pi_q \operatorname{Id} \right)}_{\to 0 \text{ as } q \to \infty}, \qquad \operatorname{div} u_q = 0$$
$$\partial_t \left(\frac{|u_q|^2}{2} \right) + \operatorname{div} \left(\left(\frac{|u_q|^2}{2} + p_q \right) u_q \right) \leq \underbrace{\operatorname{div}_{t,x} \Phi_q}_{\to 0 \text{ as } q \to \infty}$$

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Heuristics: 0

∂+

• Let $\lambda_a = a^{(b^q)}$ quantify the inverse of the diameter of an oscillation, and $\lambda_a r_a$ for $r_q \ll 1$ quantify the inverse of the distance between oscillations $(\textcircled{o}] \xrightarrow{-1} \lambda_{q}^{-1}$ $\int \xrightarrow{-1} \lambda_{q}^{-1} r_{q}^{-1}$

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- Let $\lambda_q \approx a^{(b^q)}$ quantify the inverse of the width of an oscillation, and $\lambda_q r_q$ for $r_q \ll 1$ describe the inverse of the distance between oscillations
- $u_q = \sum_{q' \leq q} w_{q'}$, where $w_{q'}$ has oscillations of size λ_q^{-1} , distance $(\lambda_q r_q)^{-1}$ between oscillations, and frequency support $[\lambda_q r_q, \lambda_q]$
- \circ "Wavelet parameter" \bar{n} quantifies the exchange between space-time and frequency support:

$$\begin{aligned} \sup_{t,x} w_{q'} \cap w_{q''} &= \emptyset \quad \text{if} \quad |q' - q''| < \bar{n} \\ \sup_{\xi} \hat{w}_{q'} \cap \hat{w}_{q''} &= \emptyset \quad \text{if} \quad |q' - q''| \geq \bar{n} \end{aligned}$$

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• Either $\mathrm{supp}\,w_q\cap\mathrm{supp}\,w_{\widetilde{q}}\equiv 0$ or $\mathrm{supp}\,\hat{w}_q\cap\mathrm{supp}\,\hat{w}_{\widetilde{q}}$





• Need to choose the support of $\mathbb{B}_{k}^{q} = \mathbb{B}_{k,\text{high}}^{q} \mathbb{B}_{k,\text{low}}^{q}$ to dodge any other bundles (for other k' or q') which have overlapping frequency support



Oscillation error and the choice of r_{q+1}

• Simple example, ignoring local energy inequality:

$$R_q - \pi_q \operatorname{Id}
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$$\int_{\mathbb{T}^3} \mathbb{B}_{q+1}^2 = 1$$
, use stationarity of $\vec{e_1} \mathbb{B}_{q+1}$, and rewrite

$$\operatorname{div} \left((R - \pi)(e_1 \otimes e_1) + w_{q+1} \otimes w_{q+1} \right) = \vec{e_1} \partial_x \left((R - \pi)(1 - \mathbb{B}_{q+1}^2) \right)$$

$$= \vec{e_1} \partial_x \left(R - \pi \right) \left(\int_{\mathbb{T}^3} -\operatorname{Id} \right) \left(\mathbb{B}_{q+1}^2 \right)$$

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• Put this vector field in divergence form and estimate in $L^{3/2}$; prefers $r_{q+1} = 1$ since $\|\mathbb{B}_{q+1}\|_{L^3}$ grows as $r_{q+1} \to 0$

• Recall that

$$w_{q+1} = \vec{e}_1 (-R + \pi)^{1/2} \underbrace{\mathbb{B}_{q+1}(x_2, x_3)}_{\text{freq's}}, \ \underbrace{\mathbb{B}_{q+1}(x_2, x_3)}_{[\lambda_{q+1}r_{q+1}, \lambda_{q+1}]},$$

and consider the (linear) Nash error

$$R_{q+1}^{\text{Nash}} = \operatorname{div}^{-1}(w_{q+1} \cdot \nabla u_q) \implies \qquad w_{q+1} \cdot \nabla u_q = \operatorname{div}\left(R_{q+1}^{\text{Nash}}\right)$$

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• Since it is linear, and \mathbb{B}_{q+1} is L^2 -normalized, it prefers *more* intermittency, so that $\|\mathbb{B}_{q+1}\|_{L^{3/2}} \to 0 \quad \text{as} \quad r_{q+1} \to 0$

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· Heuristic estimates of this choice dictate a single acceptable choice

$$r_{q+1}^{\text{Goldilocks}} = \left(rac{ ext{freq. of } R}{ ext{max. freq. of } \mathbb{B}_{q+1}}
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- Construct wild solutions to 3D Euler with well-defined helicity $\int_{\mathbb{T}^3} u \cdot (\nabla \times u)$

Thanks for your attention!